

Edge Turbulence Imaging on NSTX and Alcator C-Mod

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Goal

Understand edge turbulence by measuring its
2-D structure vs. time and comparing with theory

Outline

- Gas Puff Imaging (GPI) diagnostic
- Summary of C-Mod GPI results

Comparison of k-spectrum with DBM and BOUT

- Summary of NSTX GPI results

Radial profile during H-L mode transition

- DEGAS 2 Simulation of GPI signals
- Tentative Conclusions and Plans for 2002 run

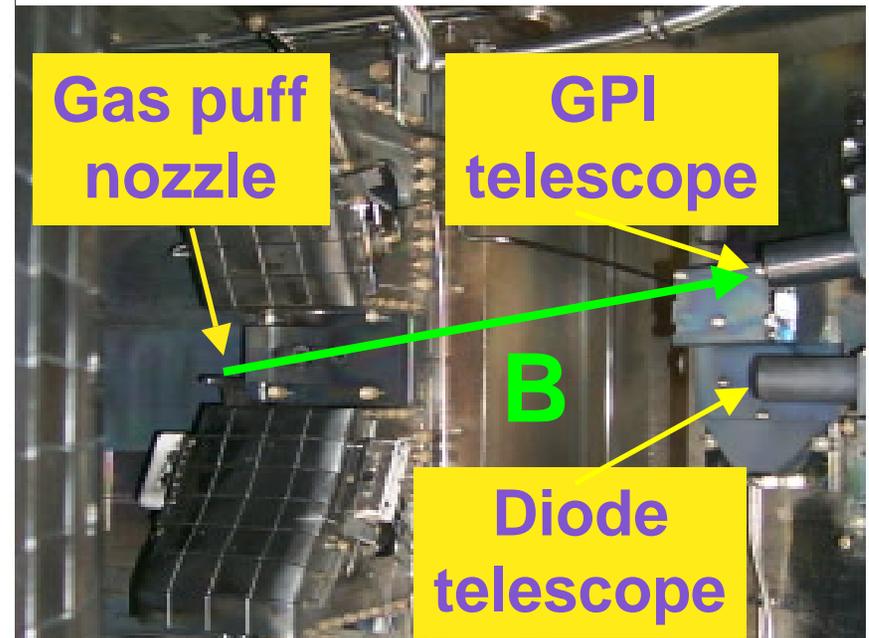
Gas Puff Imaging Diagnostic

- Gas puff imaging (GPI) telescope views neutral line emission from He or D₂ gas puff \approx along B field at the plasma edge (like BES but uses neutral gas instead of NBI)

$$S(\text{photons/cm}^3) = n_o f(n_e, T_e) A$$

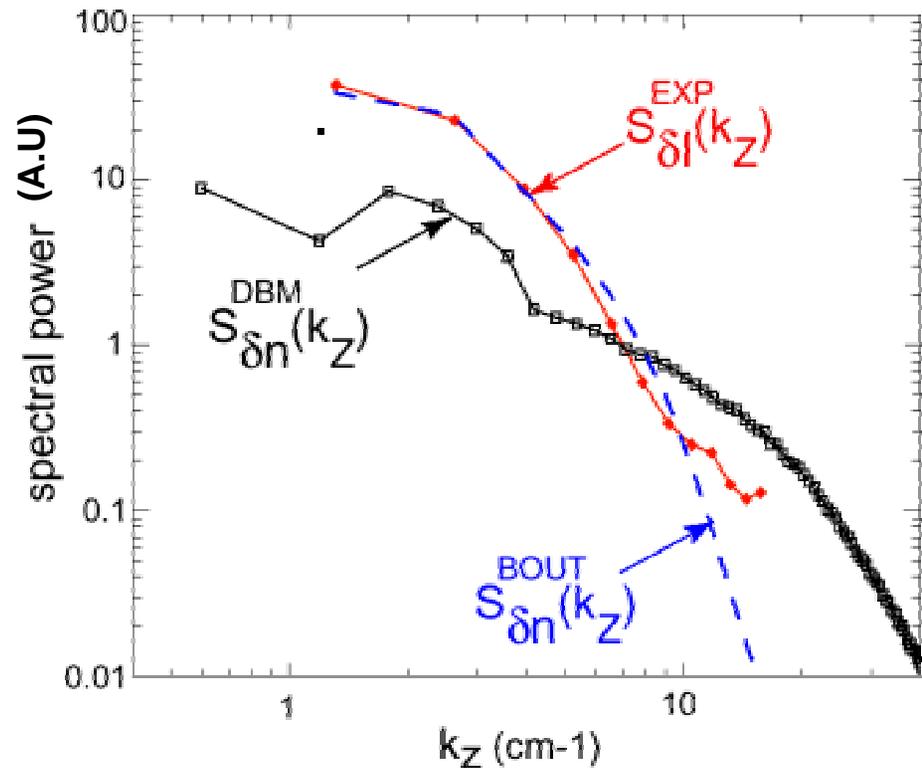
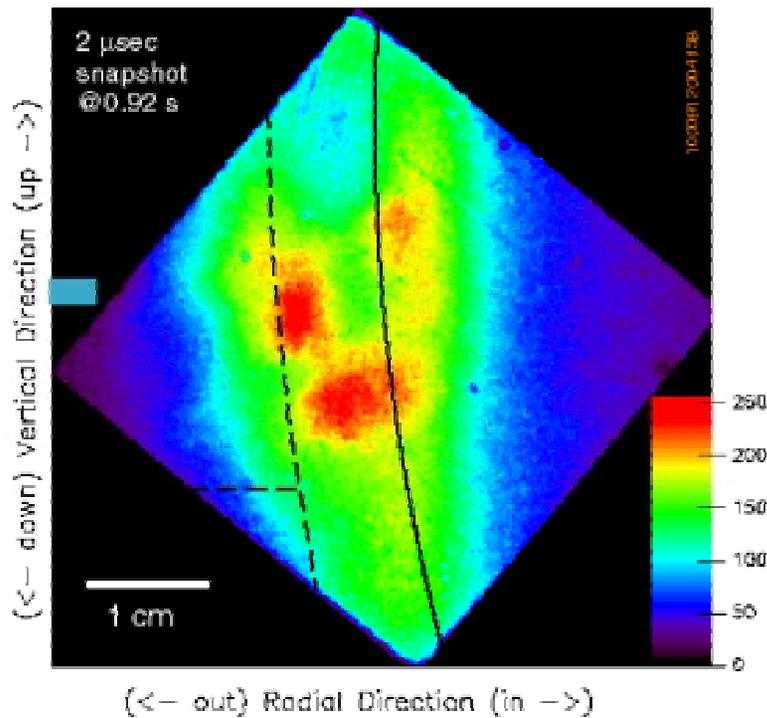
where the radiative decay rate is $A \gg 10^7 \text{ sec}^{-1}$ for these lines.

- Space and time variation of neutral light emission is measured with fast-fast gated cameras and PMs or PDs on discrete chords to determine edge turbulence structure (assumes $k_{||} \ll k_{\perp}$)
- Gas puff changes plasma density by $\approx 1\%$ in C-Mod and $\leq 10\%$ in NSTX, but this does probably not to perturb the edge turbulence significantly



Summary of C-Mod GPI Results

- Edge turbulence imaged over ≈ 6 cm poloidally x 3 cm radially at outer midplane
- Frequency spectra and fluctuation levels agree well with Langmuir probes
- Poloidal k -spectra from GPI compared with DBM code and BOUT for one case



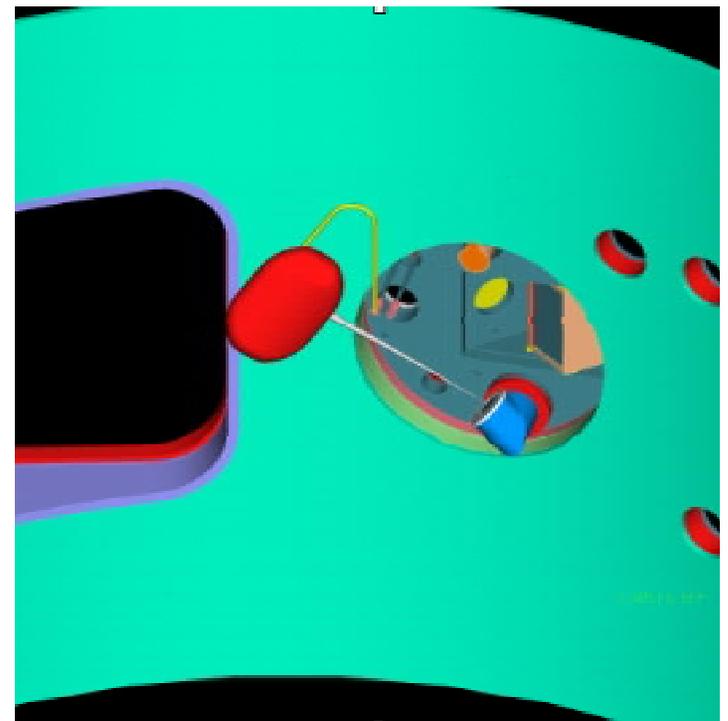
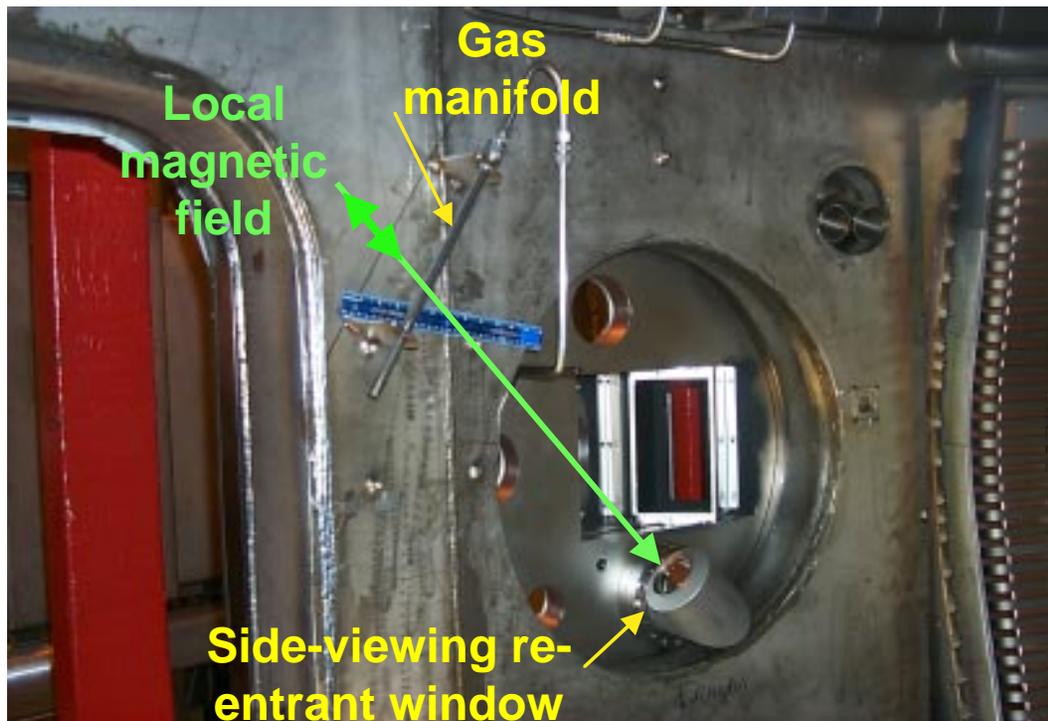
Videos of Edge Turbulence in C-Mod

- Taken at 2 μ sec/frame at 250,000 frames/sec with the 64 x 64 pixel camera made by Princeton Scientific Instruments (Model PSI-3)
- Usually appears to be “blobs” (local maxima) moving radially and poloidally through the edge and SOL

(C-mod MPEGS here)

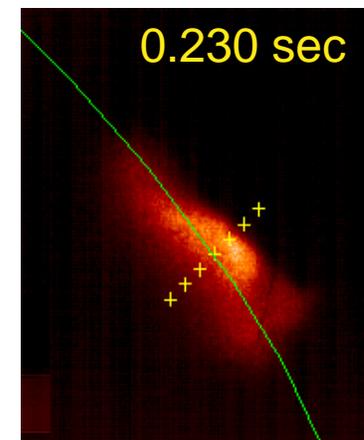
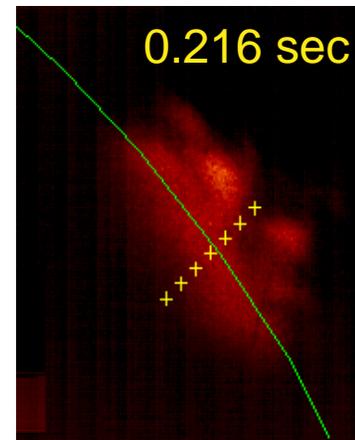
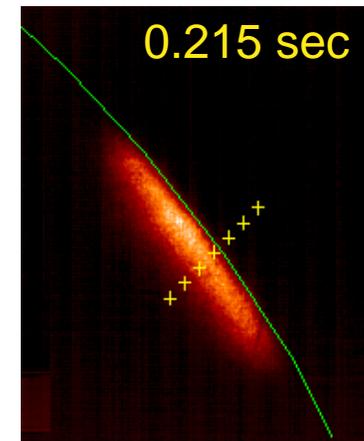
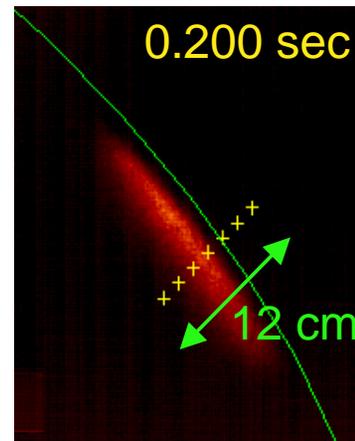
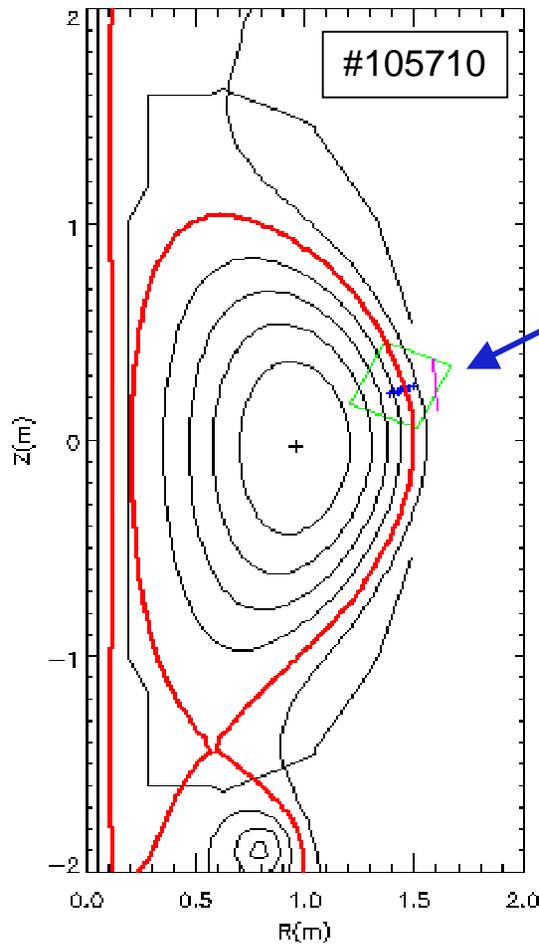
GPI Diagnostic Set-up in NSTX

- Similar to C-Mod system but using re-entrant port instead of coherent fiberbundle and elongated gas manifold instead of single-point gas nozzle
- Generally have used He puffs in NSTX and Deuterium puffs in C-Mod (although results from both are similar in each machine)



Location of GPI View (2001 Run)

- LANL Kodak camera views ≈ 30 cm x 30 cm area just above outer midplane
- Fast chords view 2 cm diameter “spots” in 7 channel radial array with PM tubes (bandwidth ≈ 100 kHz, digitized at 500 kHz)



Typical GPI Videos for NSTX

- Made using LANL Kodak camera at 10 μ sec/frame and 1000 frames/sec
- Grossly undersamples turbulence of autocorrelation time \approx 10-100 μ sec

NOTE: IMAGE ROTATED \approx 90° IN THESE MPEGS

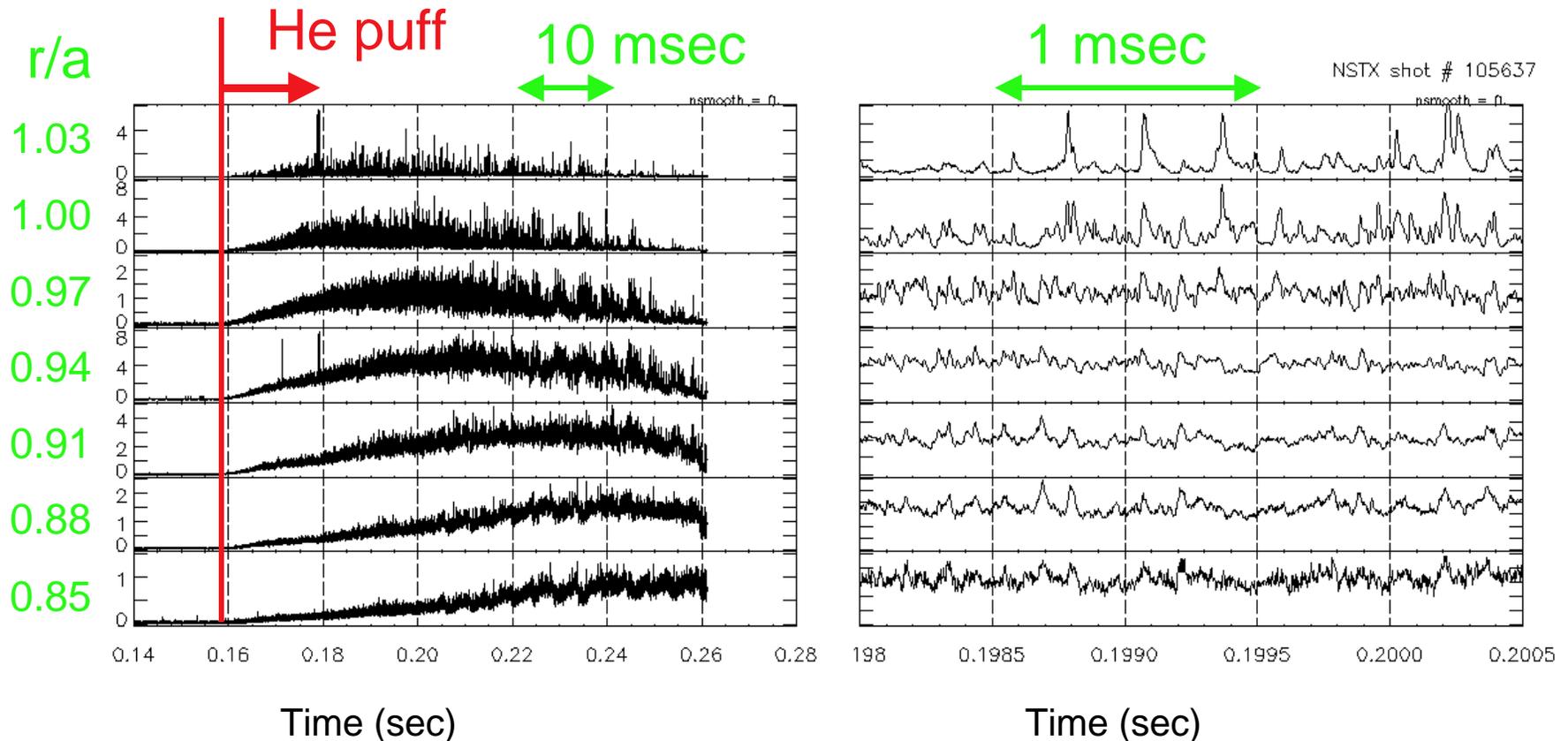
#105627
(0.7 MA, 3 kG)
density scan

#105710
(0.7 MA, 4.5 kG)
H-L transition

#105711
(0.7 MA, 4.5 kG)
no-H mode

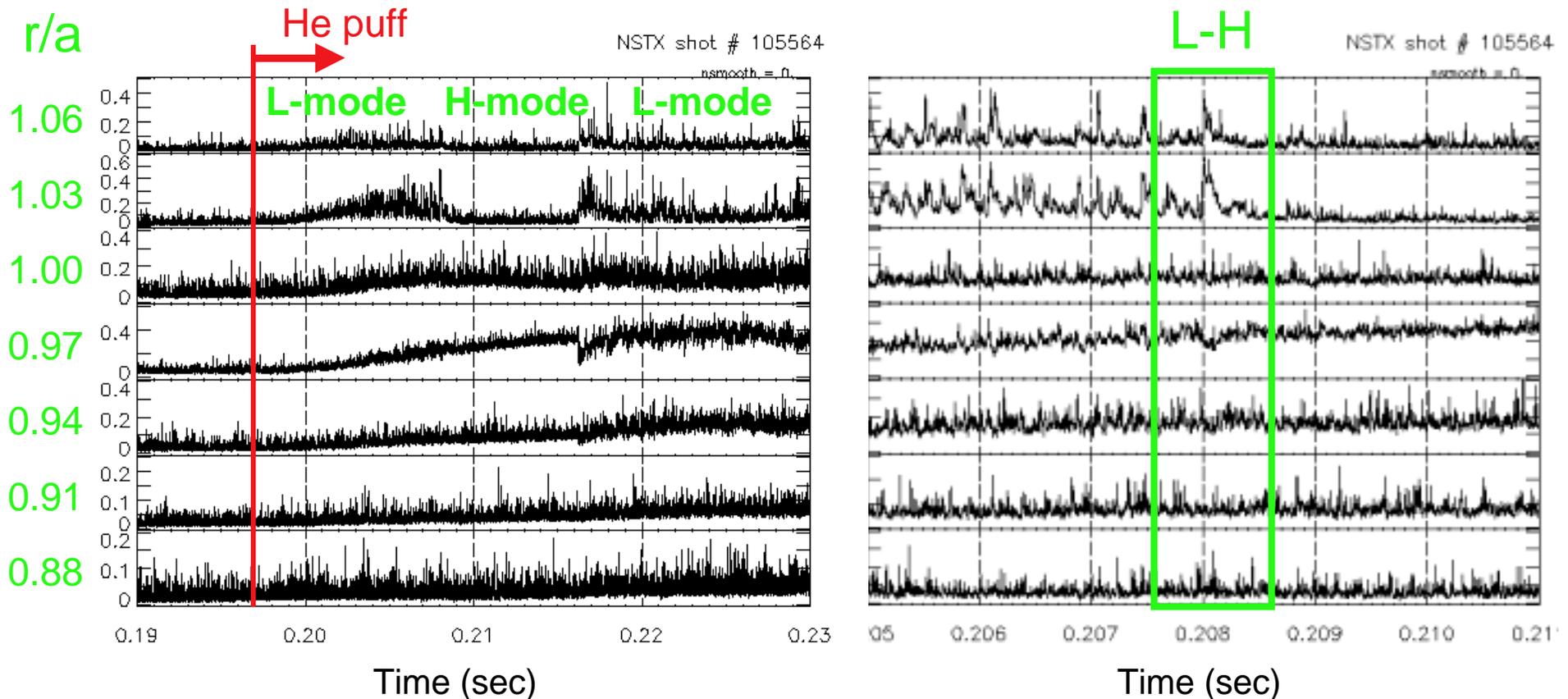
Typical Signals from Fast GPI Chords

- Signals from 7 radial chords digitized for 0.128 msec @ 500 kHz
- Near outer wall see “intermittant” fluctuations with ~ 100% modulation
- Nearer center, see “Gaussian” fluctuations with \approx 20% modulation



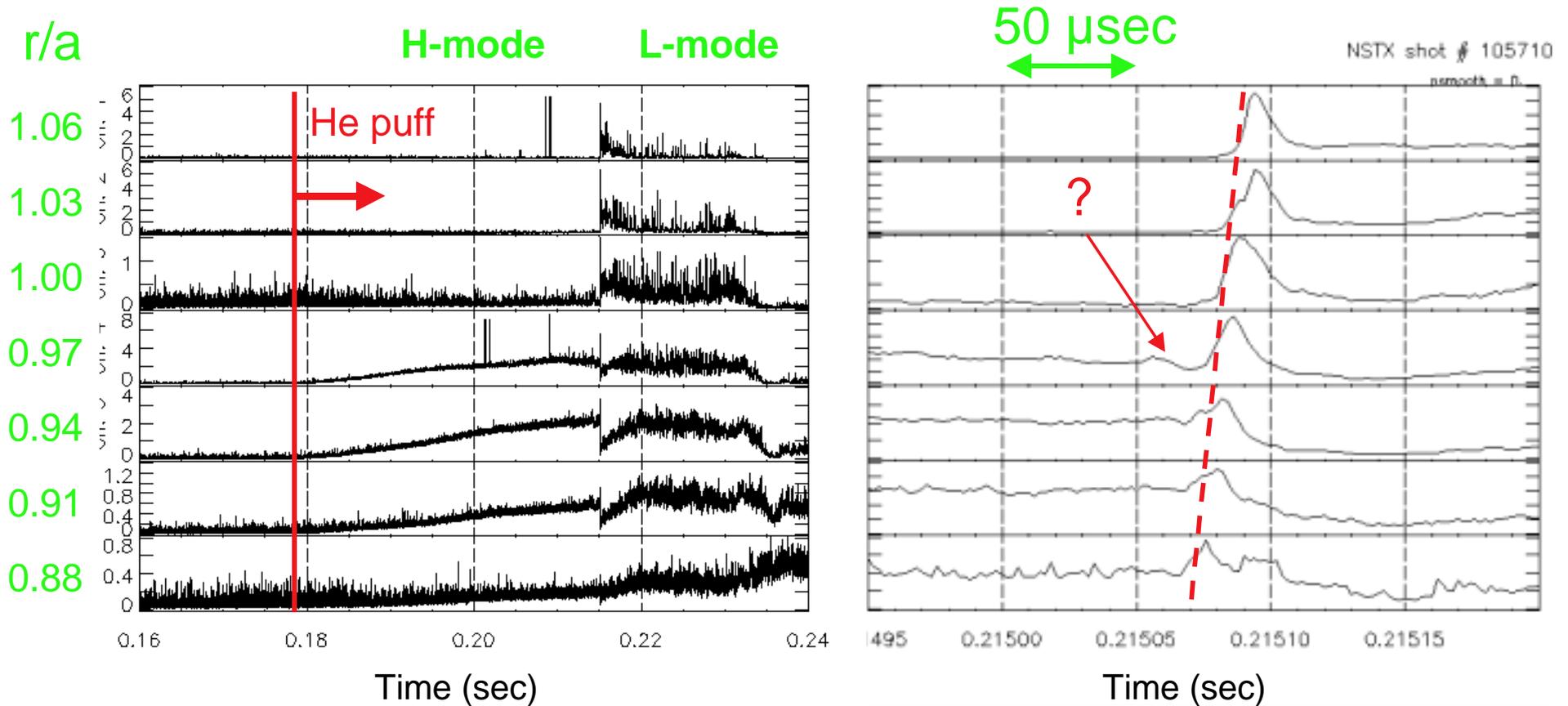
Edge Turbulence Signals at L-H Transition

- L-H transition occurs within ≈ 1 msec on most channels
- Perhaps preceded by large “bursts” in outermost channels



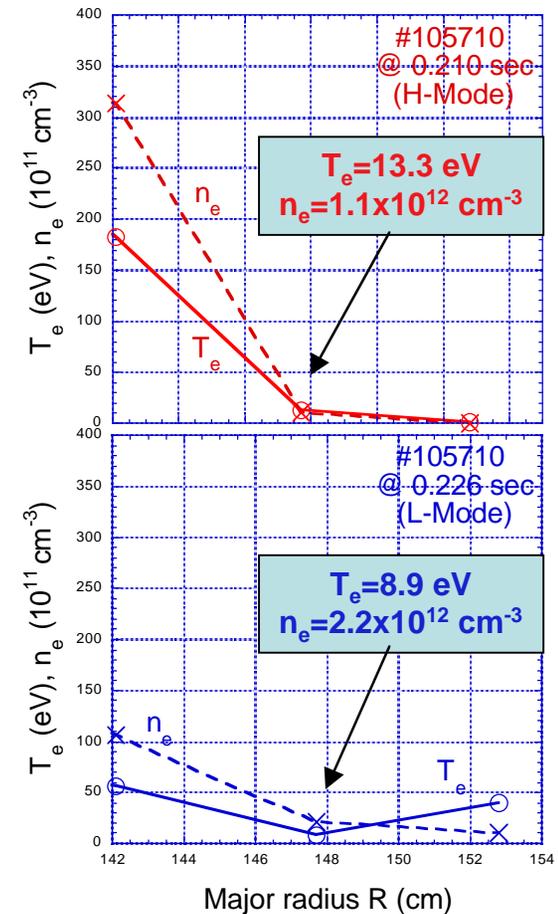
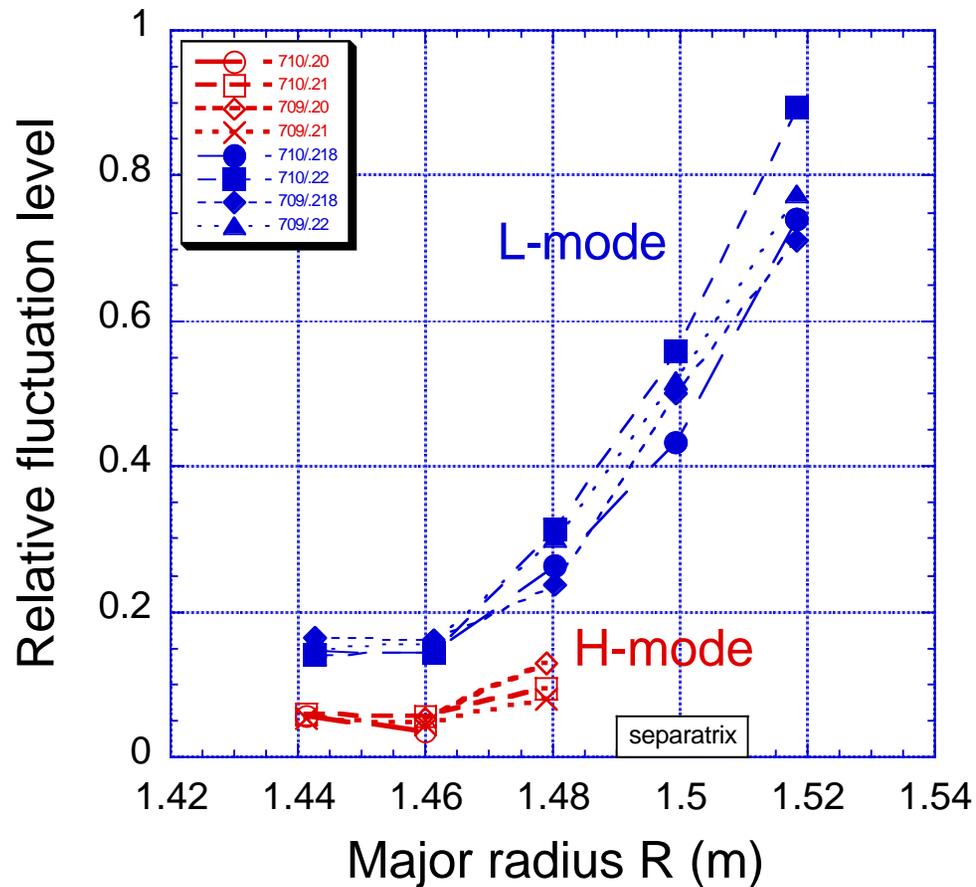
Edge Turbulence Signals at H-L Transition

- H-L transitions occurs within $\approx 20 \mu\text{sec}$ in most channels
- Transition seems to propagate outward at $\approx 10^6 \text{ cm/sec}$
- Transition seems to start with coherent mode (ELM ?)



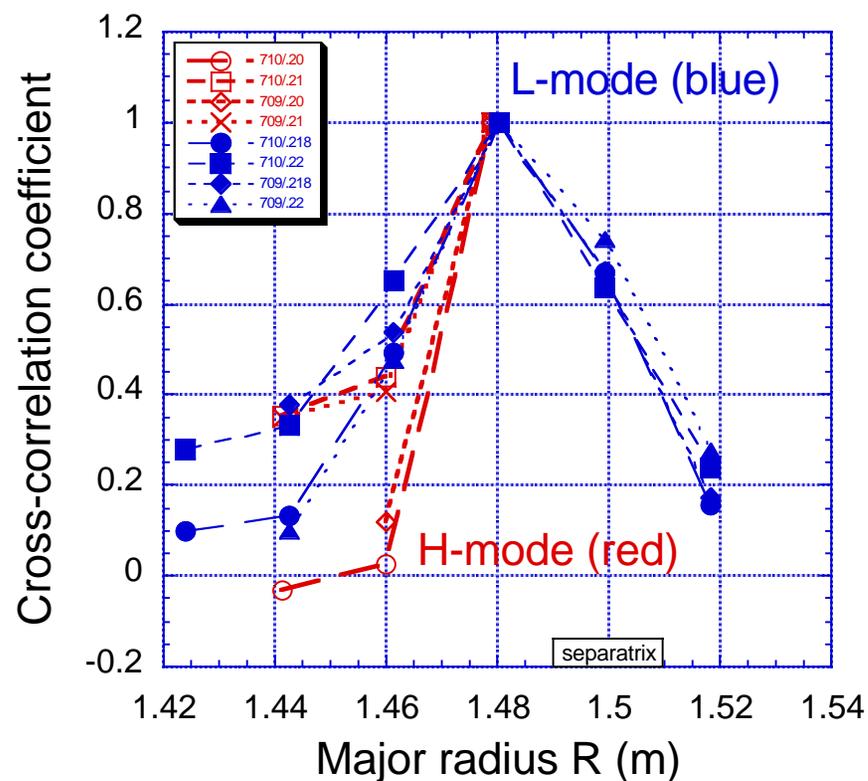
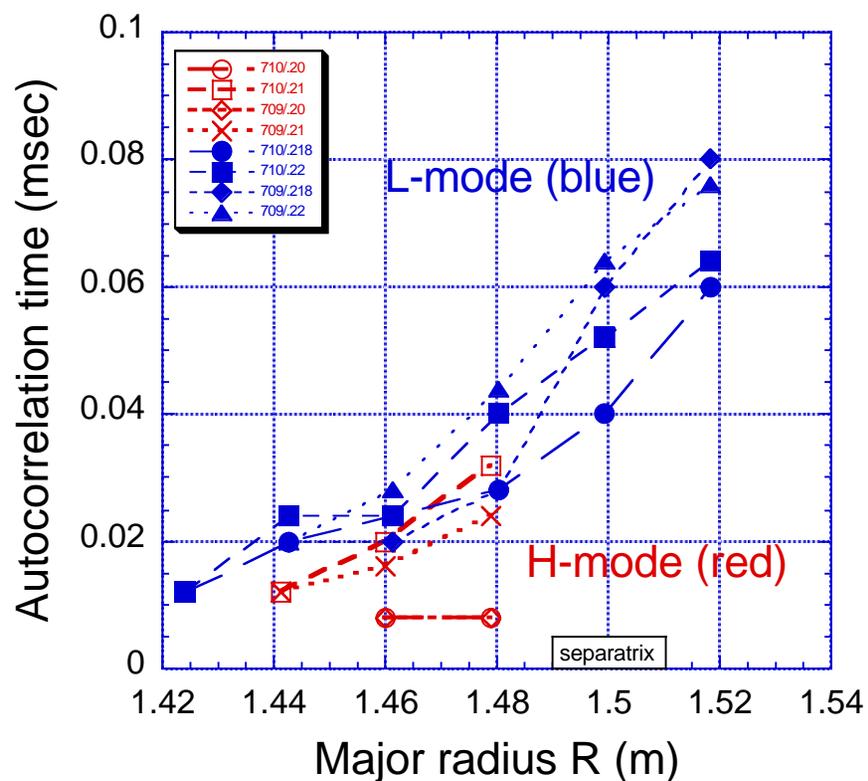
Fluctuation Profiles in H-mode vs. L-mode

- Fluctuation level seems to be lower in H-mode (but non-zero) in #105710 etc, as expected from appearance of “brain-dead” H-mode emission in images
- Only one Thomson Scattering point within range of GPI data from ‘01



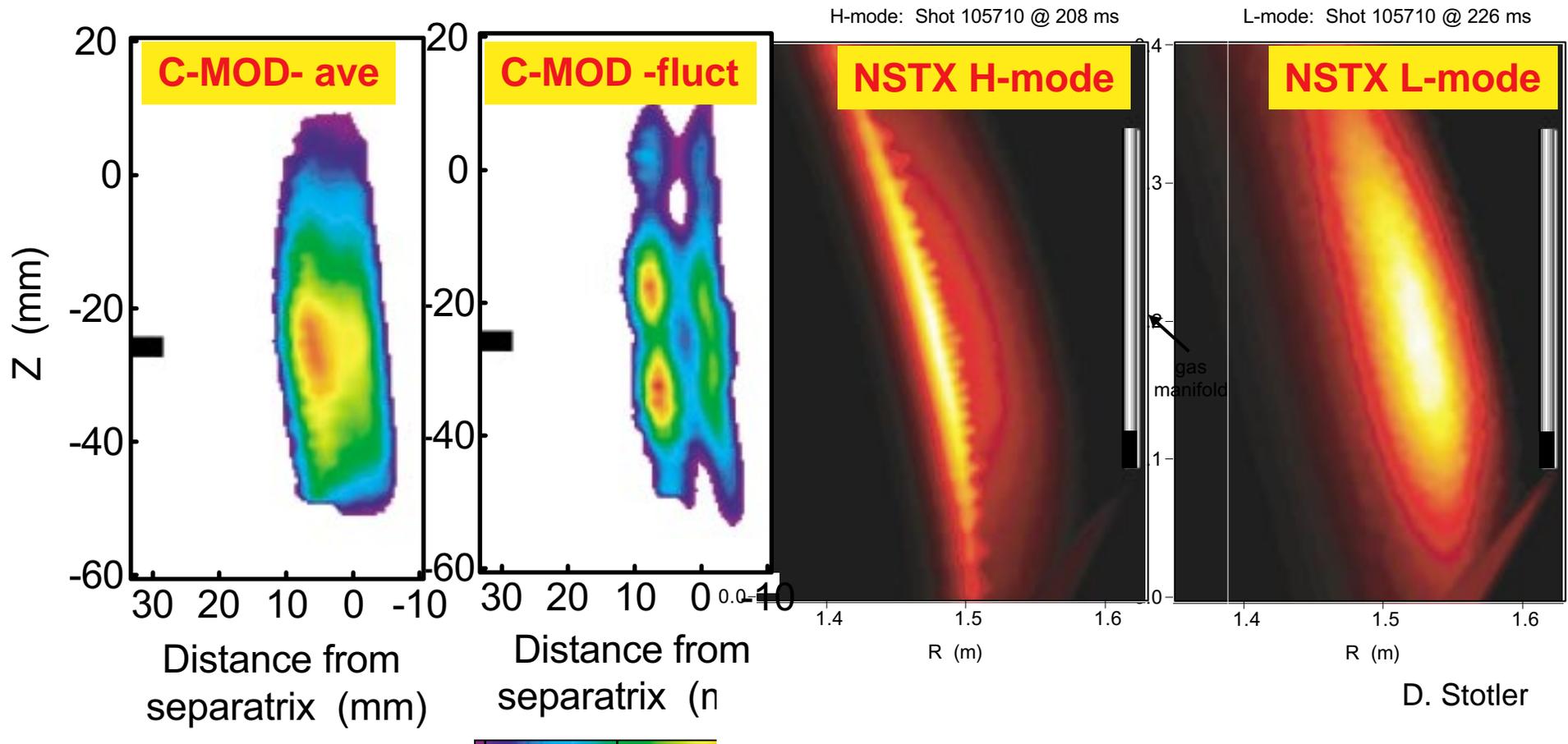
Correlations in H-mode vs. L-mode

- Correlation lengths and autocorrelation times in “late-H” (0.21 sec) similar to “L”, with cross-correlations of ≈ 0.4 implying fluctuation level is non-zero
- But correlation lengths and autocorrelation times in “early-H” (0.20 sec) lower !



DEGAS 2 Simulation of GPI

- Average D_α or He (587.6 nm) emission cloud simulated given edge profiles
- Near center of GPI cloud, $S \propto n_e^{0.5} T_e^{0.5}$ for C-Mod and $S \propto n_e^{0.7} T_e^{0.5}$ for NSTX
- H-mode emission profile is narrower simply due to narrower edge profiles



Tentative Conclusions and Plans for 2002

- Edge turbulence in GPI looks similar to that seen elsewhere
- Initial comparisons with simulation / theory are encouraging
- Interesting radial profile changes during H-L transition in NSTX
- 2-color GPI imaging in Alcator C-Mod using two Xybions
- Make more ultra-fast videos using 28 frame PSI-4 camera
- Ask theorists to run simulations for more cases
- Feed simulation into DEGAS 2 and compare with GPI data